

Water in Star-Forming Regions with Herschel (WISH): Intermediate Mass Protostar – NGC 7129



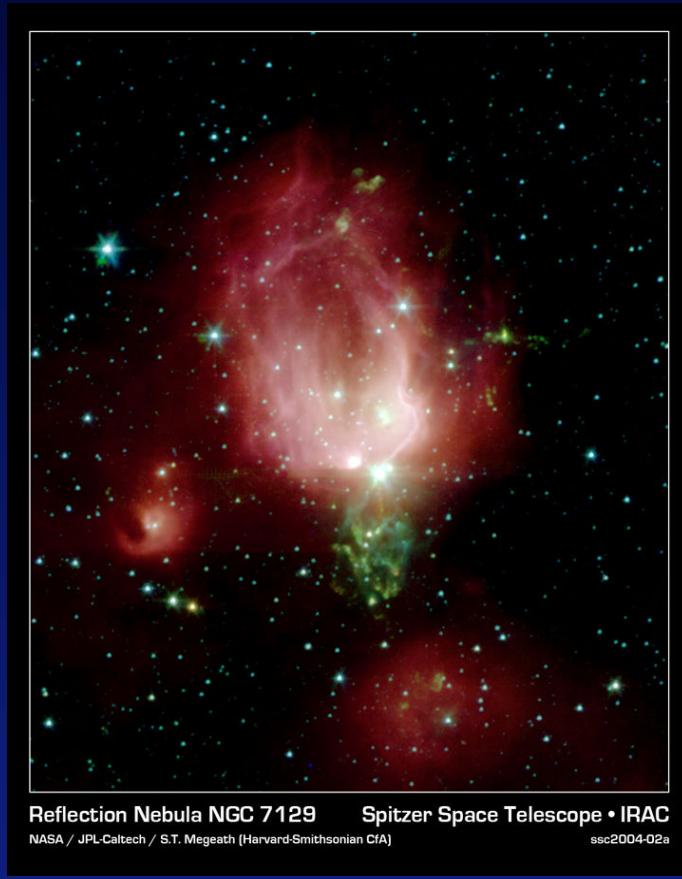
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The Herschel HIFI spectrograph is proving to be an excellent instrument for the detection of water in star-forming regions. As part of the WISH Key Program, we are undertaking a study of water in the envelope around six intermediate mass protostars: NGC 7129 FIRS 2, L1641 S3 MMS1, NGC 2071, Vela IRS 17 & 19, and AFGL 490.

The first results, toward NGC 7129 FIRS 2 (see Figure 1), have been completed during PSP and will be published in the A&A HIFI Special Edition (Johnstone et al. 2010). The water and CO spectra reveal broad and narrow components (Figures 2-4) connected to both the free-falling envelope (narrow) and the known energetic outflow (broad) around this source.

Integrated intensity fits to the free-falling envelope component require an outer envelope total water abundance of $\sim 10^{-7}$, although the precise line profiles are not well reproduced. The broad component of water emission appears related to the high- J CO lines observed with PACS (Fich et al. 2010, A&A, 518, L86) which require a very high temperature ($T \sim 1000\text{K}$) and density ($n \sim 10^8 \text{ cm}^{-3}$) shock region.

Figure 1: NGC 7129 (Watering the Rose)



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Figure 2: Selection of H_2O Spectra

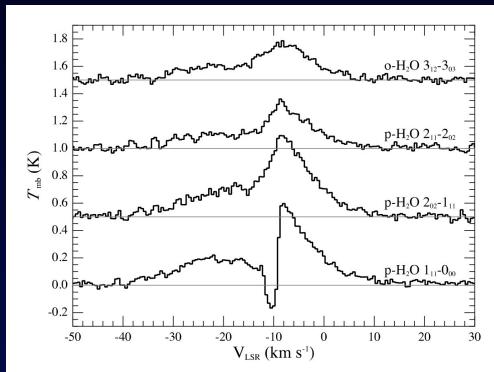


Figure 3: CO and H_2O Compared

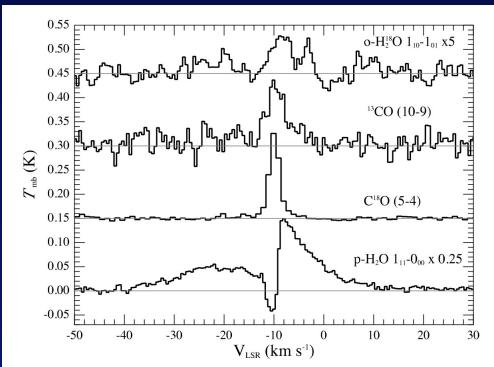


Figure 4: H_2O 2_0 - 1_1 Decomposed

